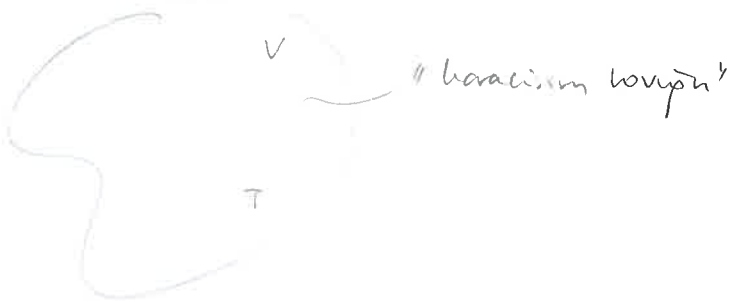


BÖLÜM 12

BOSE SİSTEMLERİ

12.1. Fotonlar



Elektromag. alan $\rightarrow$ herbiri farklı frekans ile salınan n salınan

$$\left(n + \frac{1}{2}\right) \hbar \omega \quad n = 0, 1, 2, 3, \dots$$

Fotonlar EMA'nın kuantumu.

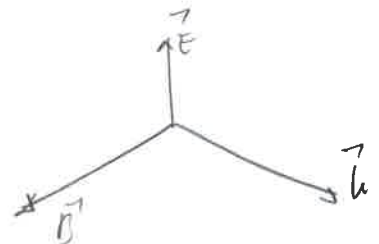
- $\hookrightarrow$ Q. Teorine göre t_0 spinli kütleli bosonlar.
- $\hookrightarrow$ kütleli fotonların c hızıyla "serbest uzayda" hareket ettiğini gösterir.
- $\hookrightarrow$ spinli sadece iki bağımsız yönelim olabilir; paralel anti paralel $\uparrow\uparrow, \downarrow\downarrow$ ($\vec{p}$ yönünde)

Foton'un belirli bir spin durumunda bulunması ya sağ elli ya da sol elli dairesel polarize bir dalganın EMA karşılığıdır. Bununla birlikte, belirli spinli iki foton durumunun süperpozisyonu ile spinin ortalaması olmayan ancak lineer polarize olan bir foton durumu elde edebiliriz. Aşağıda lineer polarize fotonlar dalgaları.

$$E_{\text{fey}} = E = \hbar \omega$$

$$\text{Momentum} = \vec{p} = \hbar \vec{k} \quad |\vec{k}| = \frac{\omega}{c}$$

$$\text{pol. vektör} = \vec{e} \quad |\vec{e}| = 1 \quad \vec{k} \cdot \vec{e} = 0$$



bir kuantede, $\vec{h} = \frac{2\pi}{L} \vec{n}$ $n = 0, \pm 1, \pm 2, \dots$

izin verilen mom. durumların sayı $k = k + dk$

$$\frac{V}{(2\pi)^3} 4\pi k^2 dk$$

atomlar foto yayılgayış seğımlarından toplam foto sayı bulunur.

$$\bar{E} \{ n_{\vec{k}, \epsilon} \} = \sum_{\vec{k}, \epsilon} \hbar \omega n_{\vec{k}, \epsilon} \quad \omega = c |\vec{k}|$$

$$n_{\vec{k}, \epsilon} = 0, 1, 2, \dots$$

sayı bulun eldederiz.

$$Q = \sum_{\{n_{\vec{k}, \epsilon}\}} e^{-\beta E \{ n_{\vec{k}, \epsilon} \}} = \prod_{\vec{k}, \epsilon} \sum_{n=0}^{\infty} e^{-\beta \hbar \omega n} = \prod_{\vec{k}, \epsilon} \frac{1}{1 - e^{-\beta \hbar \omega}}$$

$$Q = \prod_{\vec{k}, \epsilon} \frac{1}{1 - e^{-\beta \hbar \omega}}$$

$$\ln Q = - \sum_{\vec{k}, \epsilon} \ln (1 - e^{-\beta \hbar \omega}) = - 2 \sum_{\vec{k}} \ln (1 - e^{-\beta \hbar \omega})$$

$$\langle n_{\vec{k}} \rangle = - \frac{1}{\beta} \frac{\partial}{\partial (\hbar \omega)} \ln Q = \frac{2}{e^{\beta \hbar \omega} - 1}$$

$$U = - \frac{\partial}{\partial \beta} \ln Q = \sum \hbar \omega \langle n_{\vec{k}} \rangle$$

$$P = \frac{1}{\beta} \frac{\partial}{\partial V} \ln Q \quad \ln Q = - 2 \sum_{\vec{k}} \ln \left(1 - e^{-\beta \hbar c \frac{2\pi}{V^{1/3}} |\vec{n}|} \right)$$

$$PV = \frac{1}{3} U$$

$V \rightarrow \infty$ limitinde, $\sum_k \rightarrow \frac{V}{8\pi^2} \int d^3 k$

$$\frac{U}{V} = \frac{\hbar}{\pi^2 c^3} \int_0^\infty d\omega \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

$$U = \frac{2V}{8\pi^2} \int d^3 k 4\pi \hbar^3 \frac{\hbar c k}{e^{\beta \hbar c k} - 1}$$

$$= \int_0^\infty d\omega u(\omega, T) \quad u(\omega, T) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

Planck ıřıma yasan.

$$\int_0^\infty dx \frac{x^3}{e^x - 1} = \frac{\pi^4}{15}$$

$$\frac{U}{V} = \frac{\pi^2}{15} \frac{(kT)^4}{(\hbar c)^3}$$

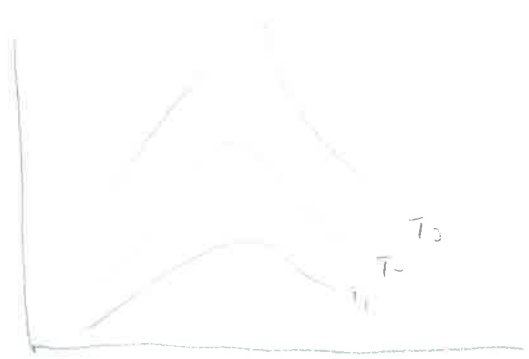
$$C_V = \frac{1}{V} \left(\frac{\partial U}{\partial T} \right)_V = \frac{4\pi^2 \hbar^4 T^3}{15 (\hbar c)^3}$$

$T \rightarrow \infty$ gide ilen bounded
segr.
kavitedeki fot. sayı
bounded olmaktadır

hıřle bir kavite re. lık deıerıer olarka
bu yasa test edilebilir. Birin den bapın bıkın remade ııııı
mıhtakı ;

$$I(\omega, T) = \frac{c}{2} \int \frac{d\Omega}{4\pi} u(\omega, T) \omega, \theta = \frac{c}{4} u(\omega, T)$$

$$I(T) \propto T^4$$



$T_1 < T_2 < \dots$

12.7. Phononen:

$$\mathcal{H} = T + V$$

N atom var $\Rightarrow 3N$ normal mod. $\omega_1, \omega_2, \dots, \omega_{3N}$

$$E\{n_i\} = \sum_{i=1}^{3N} n_i \hbar \omega_i \quad Q = \sum_{\{n_i\}} e^{-\beta E\{n_i\}} = \prod_{i=1}^{3N} \frac{1}{1 - e^{-\beta \hbar \omega_i}}$$

$$\ln Q = - \sum_{i=1}^{3N} \ln(1 - e^{-\beta \hbar \omega_i})$$

$$\langle n_i \rangle = - \frac{1}{\beta} \frac{\partial}{\partial (\hbar \omega_i)} \ln Q = \frac{1}{e^{\beta \hbar \omega_i} - 1}$$

$$U = - \partial_{\beta} \ln Q = \sum_{i=1}^{3N} \hbar \omega_i \langle n_i \rangle = \sum \frac{\hbar \omega_i}{e^{\beta \hbar \omega_i} - 1}$$

$V \neq 0$ then $\frac{U}{V} = \frac{3}{2\pi^2 c^3} \int_0^{\omega_m} d\omega \omega^3 \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$

$\omega \rightarrow \omega + d\omega$ annulieren
normal lin'leir summe

$$= \int \omega d\omega = \frac{3}{8\pi^2} 4\pi \hbar^3 d\omega$$

normale polarisierbarkeit

$$= V \frac{3\omega^3}{2\pi^2 c^3} d\omega$$

$$\int_0^{\omega_m} f(\omega) d\omega = 3N$$

$$U = V/N$$

$$\omega_m = c \left(\frac{6\pi^2}{V} \right)^{1/3}$$

$$\lambda_m = \frac{2\pi c}{\omega_m} = \left(\frac{4}{3} \pi V \right)^{1/3}$$

praktisch von nennend.

$$\frac{U}{N} = \frac{g(h\nu)^4}{(h\nu_m)^3} \int_0^{\beta h\nu_m} dt \frac{t^3}{e^t - 1}$$

$$D(x) \equiv \frac{3}{x^3} \int_0^x dt \frac{t^3}{e^t - 1} = \begin{cases} 1 - \frac{3}{8}x + \frac{1}{20}x^2 + \dots & x \ll 1 \\ \frac{\pi^4}{15x^3} + O(e^{-x}) & x \gg 1 \end{cases}$$

Debye fun.

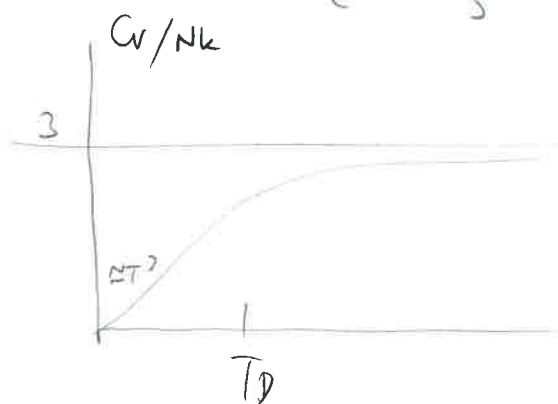
$$T_D \equiv h\nu_m = h \left(\frac{6N}{V} \right)^{1/3}$$

$$\forall \frac{U}{N} = 3kT D(\lambda) = \begin{cases} 3kT \left(1 - \frac{3T_D}{8T} + \dots \right) & T \gg T_D \\ 3kT \left(\frac{\pi^4}{15} \left(\frac{T}{T_D} \right)^3 + \dots \right) & T \ll T_D \end{cases}$$

$$\lambda = T_D / T$$

$$\frac{C_V}{Nk} = 3D(\lambda) + 3T \frac{dD(\lambda)}{dT} = 3 \left[4D(\lambda) - \frac{3\lambda}{e^{\lambda} - 1} \right]$$

$$= \begin{cases} 3 \left[1 - \frac{1}{2} \left(\frac{T_D}{T} \right)^2 + \dots \right] & T \gg T_D \\ \frac{17\pi^4}{15} \left(\frac{T}{T_D} \right)^3 + \dots & T \ll T_D \end{cases}$$



12.3. Bose-Einstein Yöğünleşme:

$$\frac{1}{V} = \frac{1}{\lambda^3} g_{3/2}(z) + \frac{1}{V} \frac{z}{1-z} \quad v = \frac{V}{N} \quad \lambda = \sqrt{2\pi\hbar^2/mkT}$$

(i) $z = e^{\beta\mu}$

$0 \leq z \leq 1$

$$G_n(z) = \int_0^\infty \frac{x^{n-1}}{z^{-1}e^x - 1} dx$$

$$\lim_{z \rightarrow 0} G_n(z) = \int_0^\infty z x^{n-1} e^{-x} dx = \Gamma(n)$$

$$G_n(z) = \Gamma(n) g_n(z) \Rightarrow g_n(z) = \frac{1}{\Gamma(n)} G_n(z)$$

$$g_n(z) = \frac{1}{\Gamma(n)} \int_0^\infty \frac{x^{n-1}}{z^{-1}e^x - 1} dx$$

örneğin $n=1$ için,

$$g_1(z) = \frac{1}{\Gamma(1)} \int_0^\infty \frac{x^{0}}{z^{-1}e^x - 1} dx$$

$$= \frac{1}{\Gamma(1)} \int_0^\infty x^{0} z e^{-x} (1 + z e^{-x} + z^2 e^{-2x} + \dots) dx$$

$$= \frac{1}{\Gamma(1)} \int_0^\infty x^{0} \sum_{l=1}^\infty (z e^{-x})^l dx$$

$$= \frac{1}{\Gamma(1)} \sum_{l=1}^\infty z^l \int_0^\infty x^{0} e^{-xl} dx$$

$\times l = 1$

$$= \frac{1}{\Gamma(n)} \sum_{l=1}^{\infty} z^l \int \frac{u^{n-1}}{e^{nu}} e^{-u} \frac{du}{u}$$

$$= \frac{1}{\Gamma(n)} \sum_{l=1}^{\infty} z^l \frac{\Gamma(n)}{l^n} = \sum_{l=1}^{\infty} \frac{z^l}{l^n} = z + \frac{z^2}{2^n} + \frac{z^3}{3^n} + \dots$$

$$\zeta_n(1) = \sum_{l=1}^{\infty} \frac{1}{l^n} = \zeta(n) \quad n > 1$$

$n < 1$: nicht absum.

$$\zeta_2(1) = \zeta(2) = \frac{\pi^2}{6} \approx 1.645$$

$$\zeta_4(1) = \zeta(4) = \frac{\pi^4}{90} \approx 1.082$$

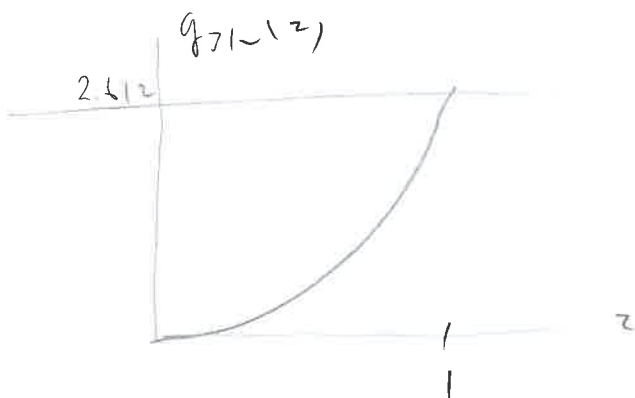
$$\zeta(6) = \frac{\pi^6}{945} \approx 1.017$$

$$\zeta(8) = \frac{\pi^8}{9450} \approx 1.004$$

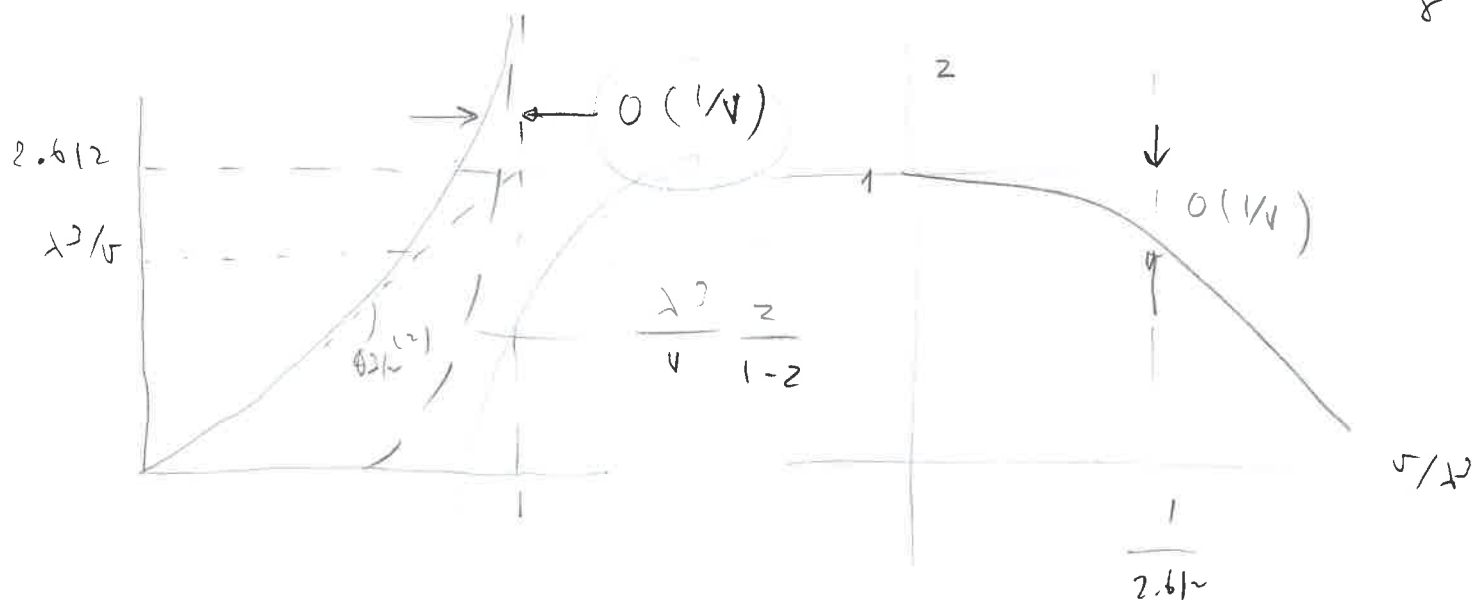
$$\zeta_{3/2}(1) = \zeta(3/2) \approx 2.612$$

$$\frac{z}{1-z} = \zeta(1)$$

$$\frac{\lambda^2 \langle n \rangle}{V} = \frac{\lambda^2}{\sigma} = \zeta_{3/2}(z) \quad \frac{\lambda^2}{\sigma} \geq \zeta_{3/2}(2)$$



z für T mit fest. v. oberst
 kommt, da dann
 größtes ist und geht.



$V \rightarrow \infty$ then O der.

$$\frac{\lambda^3}{\nu} = g_{3/2}(1) \quad (V, T)$$

$V_c \quad T_c$

$$\left(\frac{2\pi \hbar^2}{mkT_c} \right)^{3/2} = g_{3/2}(1) \quad \text{hacim abt.}$$

$$kT_c = \frac{2\pi \hbar^2 / m}{[\nu g_{3/2}(1)]^{2/3}}$$

$$\nu_c = \frac{\lambda^3}{g_{3/2}(1)}$$

$$\frac{\lambda^3}{\nu} \geq g_{3/2}(1) = 2.612$$

$$\frac{\lambda^3}{\nu} = \frac{\langle n \rangle \lambda^3}{\nu} + g_{3/2}(1) \quad z=1$$

DE yapmam

$$\langle n_0 \rangle = \frac{z}{1-z}$$

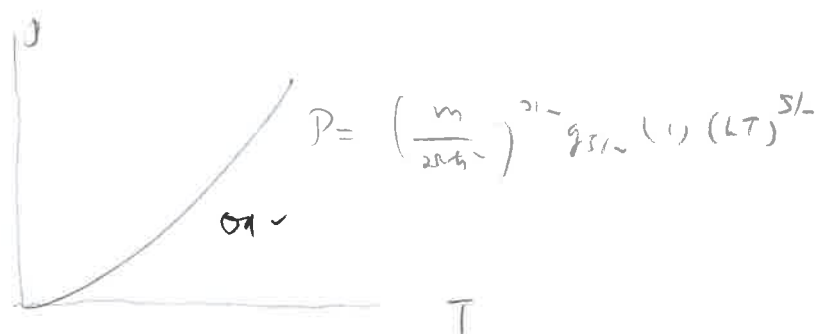
$$\frac{\lambda^3}{\nu} \langle n_0 \rangle = \frac{N \lambda^3}{\nu} - g_{3/2}(1)$$

$$\frac{\langle n_0 \rangle}{N} = 1 - \frac{\nu}{N} \frac{1}{\lambda^3} g_{3/2}(1) = 1 - \frac{g_{3/2}(1)}{g_{3/2}(z)}$$

$$\frac{P}{kT} = \begin{cases} \frac{1}{\lambda^3} g_{5/2}^{(2)} & \nu > \nu_c \\ \frac{1}{\lambda^3} g_{5/2}^{(1)} & \nu < \nu_c \end{cases}$$

$$\nu < \nu_c \quad P_0 = \frac{kT}{\lambda^3} g_{5/2}^{(2)}$$

$$\frac{dP_0}{dT} = \frac{1}{T\nu_c} \left[\frac{5}{2} kT \frac{g_{5/2}^{(1)}}{g_{3/2}^{(1)}} \right]$$



$$L = \frac{5}{2} kT \frac{g_{5/2}^{(1)}}{g_{3/2}^{(1)}}$$

$$\frac{C_v}{kT} = \begin{cases} (15/4) (5/\lambda^2) g_{5/2}^{(1)} & T < T_c \\ (15/4) (5/\lambda^2) g_{5/2}^{(2)} - (9/4) (g_{3/2}^{(2)}/g_{1/2}^{(2)}) & T > T_c \end{cases}$$

